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Shifts of the UV-visible absorption peaks of soot nanoparticles as a function of their coalescence: an atomistic study



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Among the methods used to estimate the absorption cross-section of dust particles in Earth's atmosphere, the Atomic Dynamic Point Dipole Interaction (ADPDI) model has the advantage of being sensitive to the atomistic geometry of the particles and to the chemical functions it contains [1]. However, it requires the knowledge of the atomic coordinates and polarizabilities of all atoms in the particle. Using LAMMPS, We computed the relaxed atomic coordinates for 2 onion-like carbon spherules more or less interpenetrated. As a result of the interpenetration distance, a decrease of aromatic over aliphatic C ratio is obtained. Then, We show that this causes possibly measurable shifts of the maxima of the absorption peaks of the absorption cross-section spectra, computed with ADPDI model, that increase with the degree of interpenetration.

Computational method:

Similarly to DDA [2], we can compute dipoles on each atom *i* with ADPDI model by solving the following equation system, for all *i* from 1 to *N* :

 $\vec{p}_i(\omega) = \bar{\bar{\alpha}}_i(\omega)\vec{E}_0(\vec{r}_i,\omega) + \sum \bar{\bar{\alpha}}_i(\omega)\bar{\bar{T}}(\vec{r}_i,\vec{r}_i,\omega)\vec{p}_i(\omega)$

Because in ADPDI the point dipoles are supposed to represent atoms much smaller than the discretization volumes used in the DDA [2], the i = j terms of the interaction tensor \overline{T} are assumed to simplify to :

Modelization:



TEM images at different scales of a soot nanoparticle.

$$\sum_{j=1}^{n} i(j) = j(j) = j(j)$$

Where $\overline{\overline{\alpha}}_i(\omega)$ is the atomic polarizabilities tensor, $\vec{E}_0(\vec{r}_i, \omega)$ the incident electric field applied to the particle and $\overline{T}(\vec{r}_i, \vec{r}_{i\neq i}, \omega)$ that can be computed with the double gradient of the generalized Green's function for the Helmholtz equation :

 $\bar{\bar{T}}(\vec{r}_i, \vec{r}_{j\neq i}, \omega) = -\frac{1}{\varepsilon_0} \left(\nabla_{\vec{r}_i} \otimes \nabla_{\vec{r}_j} + \frac{\omega^2}{c^2} \bar{\bar{I}} \right) \left(-\frac{e^{i\frac{\omega}{c}|\vec{r}_i - \vec{r}_j|}}{4\pi |\vec{r}_i - \vec{r}_j|} \right)$

With c the speed of light, ε_0 the vacuum permittivity and \overline{I} the identity tensor.

$$\overline{\overline{T}}(\vec{r}_i,\vec{r}_i,\omega) = i\frac{2}{3}\frac{\omega^3}{c^3}\frac{1}{4\pi\varepsilon_0}\overline{\overline{I}}$$

Once the values of the dipoles are selfconsistently computed, they can be used to compute various optical quantities of interest (such as Müller matrix or the extinction, diffusion and absorption crosssections) as in DDA [2].



Results of the relaxation:

| Parameters : | |
|---|-----|
| Interatomic potential : AIREBO [5] | |
| Minimisation algorithm : Conjugate Gradie | ent |
| Temperature : 5 K | |
| Ensemble : NVT | |
| Timestep for equilibration : 1 fs | |
| Equilibration time : 10 ps | |

m

The first peak in the RDF of the non-optimized structure is probably due to the random position of atoms in the coalescence zone. it disappears during the



Computation of the absorption cross sections:



The anisotropic atomic polarizabilities used for these computations were taken from the parameters given by Draine [6] for graphite

| inimization, when atoms | ien atoms are rearranging Maximum interpenet | | | | | |
|---|---|------|--|--|--|--|
| Molecule after equilibration phase | | | | | | |
| Total No. of atoms | 5506 | 5506 | | | | |
| Distance between the surface layer of the two particles | +3 Å | -5 Å | | | | |
| Aromatic/Aliphatic C ratio | 8.27 | 6.78 | | | | |
| 5-members rings | 594 | 474 | | | | |
| 6-members rings | 223 | 178 | | | | |
| 7-members rings | 204 | 139 | | | | |

Before coalescing, the soot grains have defects on their surface layer i.e. a random distribution of holes created by removing 10% of carbon atoms from the outer fullerene.

The Aromatic/Aliphatic C ratio decreases by 18 % as the surface layer of one particle reaches the secondary layer of the other.

| 0 | | Energy (eV) | Wavelength (nm) | MAC (m^2/g) at + 3 Å | MAC (m^2/g at - 5 Å | |
|---|------|---|--------------------|-------------------------|------------------------|--|
| 0.5 3.2 5.9 8.6 11.3 14.0 16.7 19.4 22.1 | 24.8 | 4.99 | 248.46 | 30.358 | 28.178 | |
| Energie (eV) | | 5.03 | 246.48 | 30.436 | 28.277 | |
| The first peak correspond to free-electron peak, the one that is usually measured experimentally. There is no shift visible when the surface layer of the particle are in contact but we have a shift in position and amplitude, that could be experimentally measured [7], when particles are coalescing. | | 5.05 | 245.50 | 30.455 | 28.293 | |
| | | 5.07 | 244.54 | 30.456 | 28.277 | |
| | | 5.15 | 240.74 | 30.553 | 28.143 | |
| | | 5.17 | 239.81 | 30.537 | 28.078 | |
| | | 5.23 | 237.06 | 30.355 | 27.743 | |
| | | | | | | |
| Verification of the influence of | 50 | — MAC | _AVG-5A | - MAC_AVG | 5_3A | |
| the change of geometry: Computation of the MAC with the same isotropic atomic polarizabilities for every type of atoms, thus fixing the parameters of chemical composition. No effect caused by the modification of the geometry due to the coalescence is visible. | | | | | | |
| | | 5 | | | | |
| | | 5 | | | | |
| | | 0 0.5 3.2 5.9 8.6 11.3 14.0 16.7 19.4 22.1 24.8 Energia (a)/) | | | | |
| | | | | | | |
| | | \sim $<$ | | | | |

References: Conclusion: The aim of this work was to link the morphological variations generated by the interpenetration of two soot grains to the optical response of those particles. As expected, the necking of spherules leads to an evolution in the type of 72, 333-848, **1988** carbon atom bonds, particularly in the coalescence zone. This results in a shift in the position and amplitude of the absorption peaks, which could be measured experimentally. It would therefore be possible to relate differences in structure on an atomic scale to variations in optical properties using the model developed at Besançon, which would open up the possibility of using a classification of experimental spectra to make an inverse identification.

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Acknowledgements: The calculations related to the ADPDI methods have been

performed on the super-computer facilities of the Mesocentre de calcul de Franche-Comté which is gratefully acknowledged.

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